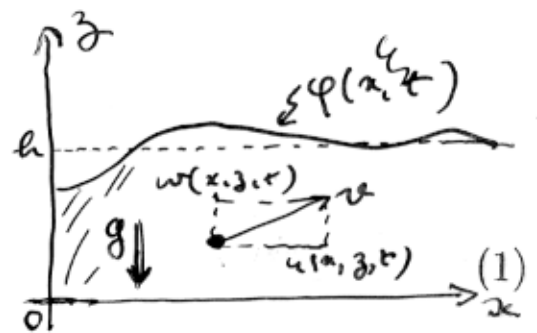


Euler equations:

$$u_t + uu_x + wu_z = -p_x$$



(1)

$$w_t + ww_x + ww_z = -p_z - g \quad (2)$$

where $p(x, z, t)$ is the pressure field and g the acceleration of gravity.

The surface is described by the function $\varphi(x, t)$ which obeys the standard continuity equation

$$\varphi_t + u\varphi_x - w = 0 \quad \text{at } z = \varphi(x, t) \quad (3)$$

The other boundary condition at the surface is

$$p(x, \varphi, t) = p_0 - \frac{T\phi_{xx}}{(1 + \phi_x^2)^{3/2}} \quad (4)$$

We take the bottom to be flat and fixed at $z = 0$ which gives the boundary condition

$$w(x, 0, t) = 0 \quad (5)$$

Linearisation: dispersion relation

$$\Omega^2 = (gk + Tk^3)\tanh kh \quad (6)$$

Green-Naghdi approximation:

$$u = u(x, t) \quad (7)$$

$$w = -zu_x(x, t) \quad (8)$$

$$\Omega^2 = \frac{ghk^2 + Thk^4}{1 + \frac{hk^2}{3}} \quad (9)$$

Multiscale analysis:

→ long waves (k small)

$$\Omega = \sqrt{gh} \left[k - \frac{1}{6} k^3 h^2 + O(k^5) \right]$$

$$v = \frac{\Omega}{k} = \sqrt{gh} \xrightarrow{x \rightarrow x - vt, t} \Omega = -\frac{1}{6} k^3 h^2 \sqrt{gh}$$

$$k = O(\epsilon) \approx \frac{\partial}{\partial x} \quad \Omega = O(\epsilon^3) = \frac{\partial}{\partial t}$$

rescale x and t to arrive at

$$\epsilon^3 \frac{\partial u}{\partial t} = \frac{1}{6} k^2 \sqrt{gh} \frac{\partial^3 u}{\partial x^3} \epsilon^3$$

then rescale u so that first non linear term is $\epsilon^3 u u_x \rightarrow$ KdV (or higher KdV)

- short waves (k large) (M. MANNING)

$$\Omega = \sqrt{\frac{3T}{h}} k + \frac{A}{k} + O\left(\frac{1}{k^3}\right)$$

$$k = O\left(\frac{1}{\epsilon}\right) = \frac{\partial}{\partial x} \quad \Omega = O(\epsilon) = \frac{\partial}{\partial t}$$

$$v = \sqrt{\frac{3T}{h}} \xrightarrow{(x,t) \rightarrow (x - vt, t)} \Omega = \frac{A}{k}$$

$$\frac{\partial^2 u}{\partial x \partial t} = -A u$$

then rescalings to obtain non linear terms of the same order.

Short-wave expansion:

Asymptotic variables

$$\zeta = (1/\epsilon)(x - v_p t) \quad (10)$$

$$\tau = \epsilon t \quad (11)$$

$$u = \epsilon^2 u_0 \quad (12)$$

$$u_{xt} = u - uu_{xx} - \frac{1}{2}u_x^2 + \frac{\lambda}{2}u_{xx}u_x^2 \quad (13)$$

Lagrangian:

$$\mathcal{L} = \frac{1}{2}u_x u_t + \frac{1}{2}u^2 + \frac{1}{2}uu_x^2 - \frac{\lambda}{24}u_x^4 \quad (14)$$

Lax pair:

$$L = \frac{\partial}{\partial x} + i\sqrt{E}F\sigma_3 + \frac{1}{2}\frac{u_{xxx}\sqrt{1-\lambda}}{F^2}\sigma_1 \quad (15)$$

$$M = -\frac{1}{2}\left(u - \frac{1}{2}\lambda u_x^2\right)\frac{u_{xxx}\sqrt{1-\lambda}}{F^2}\sigma_1 \quad (16)$$

$$\left\{ \begin{array}{l} -i\sqrt{E}\left(u - \frac{1}{2}\lambda u_x^2\right)F\sigma_3 - \frac{i}{4\sqrt{E}}\frac{1-u_{xx}}{F}\sigma_3 + \frac{1}{4\sqrt{E}}\frac{u_{xx}\sqrt{1-\lambda}}{F}\sigma_2 \end{array} \right. \quad (17)$$

where σ are the usual Pauli matrices, E the "eigenvalue" and

$$F^2 = 1 - 2u_{xx} + \lambda u_{xx}^2 \quad (18)$$

Non-trivial conserved quantities:

$$\partial_t F = \partial_x \left[\left(u - \frac{\lambda}{2}u_x^2 \right) F \right] \quad (19)$$

$$\mathcal{L} = \frac{1}{2}u_x u_t + \frac{1}{2}u^2 + \frac{1}{2}u u_x^2 - \frac{\lambda}{24}u_x^4 \quad (14)$$

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Change of function: from $u(x, t)$ to

$$g(y, t) = \frac{1}{\sqrt{1-\lambda}} \operatorname{Argtanh} \frac{u_{xx}\sqrt{1-\lambda}}{1-u_{xx}} \quad (20)$$

with

$$y = \int^x F dx \quad (21)$$

one finds that g satisfies the sinh-Gordon equation

$$\frac{\partial^2 g}{\partial y \partial t} = \frac{1}{\sqrt{1-\lambda}} \sinh \sqrt{1-\lambda} g \quad (22)$$

Lorentz invariance:

$$x \rightarrow ax, t \rightarrow t/a, u \rightarrow a^2 u \quad (23)$$

for arbitrary real a

